

Super no-scale models and moduli stability

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Introduction

- Strings unify gravity and gauge interactions at the quantum level.
- Particle physics : Start with classical 4D Minkowski space + implement perturbation theory to derive quantum dynamics.
- But from gravity point of view : Cosmological constant generated by quantum loops.
 - Except if susy : perturbative $\Lambda = 0$
 - If not susy at all : $\Lambda = \mathcal{O}(M_s^4)$
- Intermediate situation : **No-scale models** [Cremmer, Ferrara, Kounnas, Nanopoulos]
 - At tree level : Minkowski space + susy spontaneously broken
 - Potential $\mathcal{V}_{\text{tree}} \geq 0$ and admits $m_{3/2}$ as a modulus : flat direction

• Magnitude of the 1-loop effective potential $\mathcal{V}_{1\text{-loop}}$ is dictated by $m_{3/2}$.
For small $m_{3/2}$, does $\mathcal{V}_{1\text{-loop}}$ admit a small expectation value?

• **Generically, NO** : runaway behavior of $m_{3/2}$, other moduli destabilized, magnitude of $\mathcal{V}_{1\text{-loop}}$ still too large,...

• To find a loophole, we consider a context where all computations can be done explicitly, in perturbation theory :

• **Heterotic string**

• **Coordinate Dependent Compactification**

= “stringy Scherk-Schwarz mechanism”, to break spontaneously susy and gauge symmetry :

$$m_{3/2} = \frac{M_s}{R}$$

where R is the characteristic size of the compact space involved in the susy breaking

• Taking R large, to have $m_{3/2}$ and hopefully $|\mathcal{V}_{1\text{-loop}}|$ small,

\implies **Light towers of KK modes : They dominate in $\mathcal{V}_{1\text{-loop}}$.**

- Choose a point in classical moduli space $\langle G_{IJ} \rangle, \langle B_{IJ} \rangle, \langle \text{Wilson lines} \rangle, \dots$
- Suppose there are no scales between 0 and $m_{3/2}$.

———— cM_s : large Higgs, GUT or string scale

———— $m_{3/2}$: towers of KK modes of mass $\propto m_{3/2}$

———— 0 : n_B massless bosons and n_F massless fermions

$$\mathcal{V}_{1\text{-loop}} = \xi(n_F - n_B)m_{3/2}^4 + \mathcal{O}(e^{-cM_s/m_{3/2}}), \quad \xi > 0$$

- General case : Some scales may be lower than $m_{3/2}$.
- Switch on small deviations collectively denoted Y , to $\langle G_{IJ} \rangle$, $\langle B_{IJ} \rangle$, $\langle \text{Wilson lines} \rangle$:

———— cM_s : large Higgs, GUT or string scale

———— $m_{3/2}$: towers of KK modes of mass $\propto m_{3/2}$

———— YM_s : some of the $n_B + n_F$ states get a mass YM_s

———— 0

- $n_B(Y)$ and $n_F(Y)$ interpolate between different integer values, reached at distinct points in moduli space.

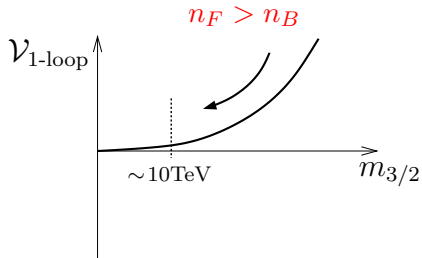
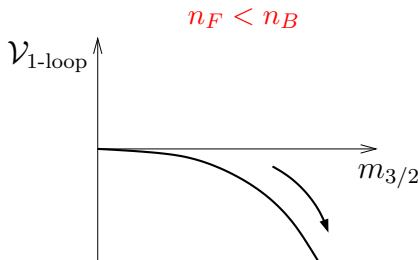
\implies Expand in Y

- For $\mathcal{N} = 4 \rightarrow \mathcal{N} = 0$

$$\mathcal{V}_{1\text{-loop}} = \xi(n_F - n_B)m_{3/2}^4 - b\tilde{\xi}m_{3/2}^2(YM_s)^2 + \dots + \mathcal{O}(e^{-cM_s/m_{3/2}})$$

- The Y 's are **Wilson lines of the non-Abelian gauge groups**.
- The b 's are their **β -function coefficients**. $\tilde{\xi} > 0$.

- **Dominant term :**



$$\langle m_{3/2} \rangle^4 \gg \Lambda_{obs}$$

$$\mathcal{V}_{1\text{-loop}} = \xi(n_F - n_B)m_{3/2}^4 - b\tilde{\xi}m_{3/2}^2(YM_s)^2 + \dots + \mathcal{O}(e^{-cM_s/m_{3/2}})$$

\implies define **“Super No-Scale Models”** in string theory by :

$$n_F = n_B \quad [\text{Kounnas, H.P.}] [\text{Abel, Dienes, Mavroudi}]$$

\implies Standard Model needs hidden sector

• **Subdominant term :**

• $b < 0 \implies Y$ stabilized at 0

• $b > 0 \implies$ **Instability** [Kounnas, H.P.]

• If there is no non-asymptotically free gauge group factor,

$$\mathcal{V}_{1\text{-loop}} = \mathcal{O}(e^{-cM_s/m_{3/2}})$$

The **Stable Super No-Scale Models** extend the notion of no-scale models to the 1-loop level :

$\mathcal{V}_{1\text{-loop}} \geq 0$ and $m_{3/2}$ is a flat direction.

- Note that in Type II and orientifold theories, there exist non-susy models with

$$\mathcal{V}_{1\text{-loop}} = 0 \quad i.e. \quad N_F = N_B \text{ are any mass level !}$$

[Kachru, Kumar, Silverstein] [Harvey]

[Shiu, Tye] [Blumenhagen, Gorlich]

[Angelantonj, Antoniadis, Forger]

[Satoh, Sugawara, Wada]

When obtained *via* spontaneous breaking of susy, they are **super no-scale models** in a “strong sense”.

- However

- $\mathcal{V}_{2\text{-loops}}$ has no reason to vanish. [Aoki, D’Hoker, Phong]

- When a perturbative heterotic dual is known, it is a conventional super no-scale models : $n_F = n_B$.

[Harvey] [Angelantonj, Antoniadis, Forger]

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Example of $\mathcal{N} = 4 \rightarrow 0$ super no-scale model

- Start from $\mathcal{N} = 4$, $E_8 \times E_8$ heterotic string on $T^2 \times T^2 \times T^2$:

$$Z = \frac{1}{\tau_2 \eta^2 \bar{\eta}^2} \frac{\Gamma(1)}{\eta^2 \bar{\eta}^2} \frac{\Gamma(2)}{\eta^2 \bar{\eta}^2} \frac{\Gamma(3)}{\eta^2 \bar{\eta}^2} (V_8 - S_8) (\bar{O}_{16} + \bar{S}_{16}) (\bar{O}_{16} + \bar{S}_{16})$$

where the left-moving worldsheet fermions contribute

$$V_8 - S_8 = \sum_{a,b} (-1)^{a+b+ab} \frac{\theta \begin{bmatrix} a \\ b \end{bmatrix}^4}{\eta^4}$$

and the lattice is modular invariant

$$\Gamma(1) = \sum_{\substack{m_1, m_2 \\ n_1, n_2}} q^{\frac{1}{2}|p_L|^2} \bar{q}^{\frac{1}{2}|p_R|^2} = \frac{\sqrt{\det G}}{\tau_2} \sum_{\substack{\tilde{m}_1, \tilde{m}_2 \\ n_1, n_2}} e^{-\frac{\pi}{\tau_2} (\tilde{m}_i + n_i \tau) (G+B)_{ij} (\tilde{m}_j + n_j \bar{\tau})}$$

- To break susy, couple lattice to the spin structure *via* a modular invariant sign, *e.g.*

$$(-1)^{\tilde{m}_1 a + n_1 b + \tilde{m}_1 n_1}$$

$$\Rightarrow \begin{cases} (-1)^{bn_1} & : \text{odd } n_1 \text{ reverse GSO} \\ m_1 + \frac{a+n_1}{2} & : \text{shifts the KK masses} \end{cases}$$

- When the first T^2 is large, all massless states have $n_1 = 0$.
 \Rightarrow The massless fermions ($a = 1$) get a KK mass $\Rightarrow n_F = 0$.
 Cannot be super no-scale.

- We need to keep some fermions massless :

$$\bar{O}_{16} + \bar{S}_{16} = \frac{1}{2} \sum_{\gamma, \delta} \frac{\bar{\theta}[\gamma]_{\delta}^8}{\bar{\eta}^8}, \quad \text{where } \gamma = 0 \Leftrightarrow \bar{O}_{16} \text{ and } \gamma = 1 \Leftrightarrow \bar{S}_{16}.$$

- Insert

$$(-1)^{\tilde{m}_1(a+\gamma+\gamma')+n_1(b+\delta+\delta')+\tilde{m}_1 n_1}$$

$$\Rightarrow \begin{cases} \text{When } \gamma + \gamma' = 0 \text{ or } 2, \text{ nothing changes.} \\ \text{When } \gamma + \gamma' = 1, \text{ roles of Bosons and Fermions reversed.} \end{cases}$$

$$\begin{aligned}
Z = & \frac{1}{\tau_2 \eta^2 \bar{\eta}^2} \frac{\Gamma(2)}{\eta^2 \bar{\eta}^2} \frac{\Gamma(3)}{\eta^2 \bar{\eta}^2} \frac{1}{\eta^2 \bar{\eta}^2} \times \\
& \left[\Gamma^{(1)}[e] \left(V_8(\bar{O}_{16} \bar{O}_{16} + \bar{S}_{16} \bar{S}_{16}) - S_8(\bar{O}_{16} \bar{S}_{16} + \bar{S}_{16} \bar{O}_{16}) \right) \right. \\
& + \Gamma^{(1)}[o] \left(V_8(\bar{O}_{16} \bar{S}_{16} + \bar{S}_{16} \bar{O}_{16}) - S_8(\bar{O}_{16} \bar{O}_{16} + \bar{S}_{16} \bar{S}_{16}) \right) \\
& + \Gamma^{(1)}[e] \left(O_8(\bar{V}_{16} \bar{C}_{16} + \bar{C}_{16} \bar{V}_{16}) - C_8(\bar{V}_{16} \bar{V}_{16} + \bar{C}_{16} \bar{C}_{16}) \right) \\
& \left. + \Gamma^{(1)}[o] \left(O_8(\bar{V}_{16} \bar{V}_{16} + \bar{C}_{16} \bar{C}_{16}) - C_8(\bar{V}_{16} \bar{C}_{16} + \bar{C}_{16} \bar{V}_{16}) \right) \right]
\end{aligned}$$

where $\Gamma^{(1)} \left[\begin{array}{l} \text{parity of winding} \\ \text{parity of momentum } 2m_1 + a \end{array} \right]$

- $m_{3/2}^2 = \frac{|U_1|^2 M_s^2}{\text{Im } T_1 \text{ Im } U_1}$ where T_1, U_1 are the Kähler and complex structure of the first T^2 .

- When $\text{Im } T_1$ is large, the gauge group is

$$U(1)^2 \times G^{(2)} \times G^{(3)} \times SO(16) \times SO(16)$$

- The massless spectrum satisfies

$$n_B = 8 \left(244 + \dim G^{(2)} + \dim G^{(3)} \right), \quad n_F = 8 \times 256.$$

12 missing bosons are obtained when T_2, U_2 and T_3, U_3 at the enhanced symmetry points

$$G^{(2)} \times G^{(3)} = SU(2)^4 \quad \text{or} \quad G^{(2)} \times G^{(3)} = SU(3) \times SU(2) \times U(1)$$

- At these points, the model develops a **super no-scale structure**.

Properties of the model

$$m_{3/2}^2 = \frac{|U_1|^2 M_s^2}{\text{Im } T_1 \text{Im } U_1}$$

- When $\text{Im } T_1 > 1, \text{Im } U_1 \sim 1$

$$m_{3/2} < M_s, \quad \mathcal{V}_{1\text{-loop}} = \mathcal{O}(e^{-cM_s/m_{3/2}}) : \quad \text{super no-scale regime}$$

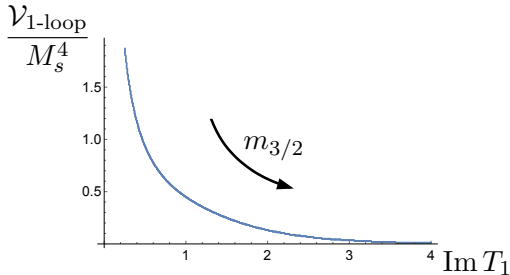
- When $\text{Im } T_1$ decreases up to ~ 1

$$m_{3/2} \sim M_s, \quad \mathcal{V}_{1\text{-loop}} \text{ is not small.}$$

Do we have an Hagedorn-like divergence of $\mathcal{V}_{1\text{-loop}}$?

- In $(-1)^{m_1 a}$ breaking, **YES** : $O_8 \bar{O}_{16} \bar{O}_{16} \implies$ Tachyons
- In $(-1)^{m_1(a+\gamma+\gamma')}$ breaking, **NO**: $O_8 \bar{V}_{16} \bar{V}_{16} \implies$ Non-level matched

- For $\text{Re } T_1 = 0$ and $U_1 = T_2 = U_2 = T_3 = U_3 = i$:



- When $\text{Im } T_1 \rightarrow 0$, the first T^2 shrinks, which is equivalent to a **dual theory in 6D, explicitly non susy**.

So, when $m_{3/2} > M_s$, the model is better interpreted as a **compactification of this non-susy theory down to 4 dimensions**.

- The model is self-dual under

$$(T_1, U_1) \longrightarrow \left(-\frac{1}{U_1}, -\frac{1}{T_1} \right)$$

So, **evolving T_1 from 0 (non susy) to $i\infty$ (super no-scale)**
 \iff **evolving U_1 from $i\infty$ (super no-scale) to 0 (non susy)**.

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Moduli deformations

- The classical moduli space of the model is

$$\frac{SO(6, 6 + 16)}{SO(6) \times SO(6 + 16)}$$

We consider the model at point where $T^2 \times T^2 \times T^2$, where the **first torus is large** and **the last two are at an enhanced symmetry point** where the model is **super no-scale**, *e.g.*

$$U(1)^2 \times SU(2)^4 \times SO(16)^2$$

We switch on an arbitrary marginal deformation of the classical theory around this point and compute $\mathcal{V}_{1\text{-loop}}$ to study the local stability :

Is the super no-scale point a minimum, maximum or saddle ?

- Compute
$$\mathcal{V}_{1\text{-loop}} = -\frac{M_s^4}{(2\pi)^4} \int_{\mathcal{F}} \frac{d^2\tau}{2\tau_2^2} Z$$

Since T^2 is large, the **winding states are super heavy** $\implies \mathcal{O}(e^{-\text{Im} T_1})$.

We keep in Z all states s_0 with 0-winding and 0-momenta along the first T^2 and take into account their **towers of KK modes**

$$(-1)^{F_0} \frac{1}{\tau_2} \sum_{m_1, m_2} (-1)^{m_1} e^{-\pi\tau_2 \frac{|U_1 m_1 - m_2 + \xi|^2}{\text{Im } T_1 \text{Im } U_1}} q^{\frac{1}{4} M_{0L}^2} \bar{q}^{\frac{1}{4} M_{0R}^2}$$

$$(-1)^{F_0} \frac{\text{Im } T_1}{\tau_2^2} \sum_{\tilde{m}_1, \tilde{m}_2} e^{-\frac{\pi \text{Im } T_1}{\tau_2 \text{Im } U_1} |\tilde{m}_1 + \frac{1}{2} + U_1 \tilde{m}_2|^2} e^{2i\pi \frac{\text{Im}[(\tilde{m}_1 + \frac{1}{2} + U_1 \tilde{m}_2)\bar{\xi}]}{\text{Im } U_1}} q^{\frac{1}{4} M_{0L}^2} \bar{q}^{\frac{1}{4} M_{0R}^2}$$

$$\implies \int_{\mathcal{F}} d^2\tau (\dots) = \int_{-1/2}^{1/2} d\tau_1 \int_0^{+\infty} d\tau_2 (\dots) + \mathcal{O}(e^{-\sqrt{\text{Im } T_1}})$$

\implies Only the level matched states contribute

Finally, $q^{\frac{1}{4} M_{0L}^2} \bar{q}^{\frac{1}{4} M_{0R}^2} = e^{-2\pi\tau_2(N_L - \frac{1}{2} + \dots)}$ with $\tau_2 \geq \text{Im } T_1$

\implies Lowest number $N_L = \frac{1}{2}$ dominates.

The **oscillators states** give $\mathcal{O}(e^{-\sqrt{\text{Im } T_1}})$ contributions.

- For small moduli deformations : **Keep only the KK towers above the initially massless states s_0**

$$\mathcal{V}_{1\text{-loop}} = -\frac{M_s^4}{(2\pi)^4} \sum_{s_0=1}^{n_B+n_F} (-1)^{F_0} \int_0^{+\infty} \frac{d\tau_2}{2\tau_2^3} \sum_{m_1, m_2} (-1)^{m_1} e^{-\pi\tau_2 M_L^2} + \mathcal{O}(e^{-c\sqrt{\text{Im} T_1}})$$

where M_L is the mass of each KK mode, deformed by the worldsheet operators

$$y_I^i \partial X_i \bar{\partial} X^I, \quad i, I = 1, \dots, 6 \quad \text{and} \quad y_I^i \partial X_i \bar{\partial} \phi^I, \quad I = 7, \dots, 22 :$$

$$M_L^2 = 2 \left(|p_L^{(1)}|^2 + \sum_{i=3}^6 (p_L^i)^2 \right)$$

$$2|p_L^{(1)}|^2 = \frac{|U_1 m_1 - m_2 + \sum_{I=3}^{22} (i \text{Im} U_1 y_I^1 - y_I^2) Q^I|^2}{\text{Im} T_1 \text{Im} U_1}$$

$$2(p_L^i)^2 = \left(\frac{m_i + \text{Re} [(i y_1^i - y_2^i) \bar{Q}^{(1)}]}{R_i} + \sum_{I=3 \neq i}^{22} y_I^i Q^I + n_i R_i \right)^2$$

- The Q 's are the charges of the KK modes, with respect to $U(1)^2 \times SU(2)^4 \times SO(16)^2$.

They can be neutral or in the Adjoint of one of the $SU(2)$'s, or Adjoint or Spinorial of one of the $SO(16)$'s.

$$\begin{aligned} \mathcal{V}_{1\text{-loop}} = & \frac{n_F - n_B}{16\pi^7} \frac{M_s^4}{(\text{Im } T_1)^2} E_{(1,0)}(U_1|3, 0) \left(1 + \sum_{j=3}^6 \left[2|Y^j|^2 - \frac{3}{2}\rho_3(Y^j)^2 - \frac{3}{2}\bar{\rho}_3(\bar{Y}^j)^2 \right] \right) \\ & - \frac{3}{16\pi^5} \frac{M_s^4}{\text{Im } T_1} E_{(1,0)}(U_1|2, 0) \left(b_{SU(2)} \sum_{i=3}^6 \left[\sum_{j=3}^6 (Y_i^j)^2 + 2|Y_i|^2 - \rho_2(Y_i)^2 - \bar{\rho}_2(\bar{Y}_i)^2 \right] \right. \\ & \left. + b_{SO(16)} \sum_{I=7}^{22} \left[\sum_{j=3}^6 (Y_I^j)^2 + 2|Y_I|^2 - \rho_2(Y_I)^2 - \bar{\rho}_2(\bar{Y}_I)^2 \right] \right) \\ & + \mathcal{O}(M_s^4 Y^3) + \mathcal{O}(M_s^4 e^{-c\sqrt{\text{Im } T_1}}) \end{aligned}$$

where
$$E_{(1,0)}(U|s, k) = \sum_{\tilde{m}_1, \tilde{m}_2} \frac{(\text{Im } U)^s}{(\tilde{m}_1 + \frac{1}{2} + \tilde{m}_2 U)^{s+k} (\tilde{m}_1 + \frac{1}{2} + \tilde{m}_2 \bar{U})^{s-k}}$$

$$\rho_s = \frac{E_{(1,0)}(U_1|s, 1)}{E_{(1,0)}(U_1|s, 0)}$$

$$\begin{aligned}
\mathcal{V}_{1\text{-loop}} = & \frac{n_F - n_B}{16\pi^7} \frac{M_s^4}{(\text{Im } T_1)^2} E_{(1,0)}(U_1|3, 0) \left(1 + \sum_{j=3}^6 \left[2|Y^j|^2 - \frac{3}{2}\rho_3(Y^j)^2 - \frac{3}{2}\bar{\rho}_3(\bar{Y}^j)^2 \right] \right) \\
& - \frac{3}{16\pi^5} \frac{M_s^4}{\text{Im } T_1} E_{(1,0)}(U_1|2, 0) \left(b_{SU(2)} \sum_{i=3}^6 \left[\sum_{j=3}^6 (Y_i^j)^2 + 2|Y_i|^2 - \rho_2(Y_i)^2 - \bar{\rho}_2(\bar{Y}_i)^2 \right] \right. \\
& \quad \left. + b_{SO(16)} \sum_{I=7}^{22} \left[\sum_{j=3}^6 (Y_I^j)^2 + 2|Y_I|^2 - \rho_2(Y_I)^2 - \bar{\rho}_2(\bar{Y}_I)^2 \right] \right) \\
& \quad + \mathcal{O}(M_s^4 Y^3) + \mathcal{O}(M_s^4 e^{-c\sqrt{\text{Im } T_1}})
\end{aligned}$$

- **1st line** $\propto m_{3/2}^4$

- It is not there in the **super no-scale models** but is there in the generic no-scale models.
- The Y^j 's are moduli that **break the $T^2 \times T^4$ factorization** and also **deform the definition of $m_{3/2}$** .
- Their mass is $\propto \frac{m_{3/2}^2}{M_s}$

$$\mathcal{V}_{1\text{-loop}} = \frac{n_F - n_B}{16\pi^7} \frac{M_s^4}{(\text{Im } T_1)^2} E_{(1,0)}(U_1|3, 0) \left(1 + \sum_{j=3}^6 \left[2|Y^j|^2 - \frac{3}{2}\rho_3(Y^j)^2 - \frac{3}{2}\bar{\rho}_3(\bar{Y}^j)^2 \right] \right) \\
- \frac{3}{16\pi^5} \frac{M_s^4}{\text{Im } T_1} E_{(1,0)}(U_1|2, 0) \left(b_{SU(2)} \sum_{i=3}^6 \left[\sum_{j=3}^6 (Y_i^j)^2 + 2|Y_i|^2 - \rho_2(Y_i)^2 - \bar{\rho}_2(\bar{Y}_i)^2 \right] \right. \\
\left. + b_{SO(16)} \sum_{I=7}^{22} \left[\sum_{j=3}^6 (Y_I^j)^2 + 2|Y_I|^2 - \rho_2(Y_I)^2 - \bar{\rho}_2(\bar{Y}_I)^2 \right] \right) \\
+ \mathcal{O}(M_s^4 Y^3) + \mathcal{O}\left(M_s^4 e^{-c\sqrt{\text{Im } T_1}}\right)$$

• **2nd and 3rd lines** $\propto m_{3/2}^2 M_s^2$

- The Y 's are the **Wilson lines** of the four $SU(2)$'s and of the two $SO(16)$ along T^6 .
- Their masses are $\propto m_{3/2}$
- $b_{SU(2)} = -\frac{8}{3} \times 2 < 0 \implies$ **Moduli stabilized at the origin.**
- $b_{SO(16)} = +\frac{8}{3} \times 2 > 0 \implies$ **Tachyonic** : They condense.
Go to **new vacuum where the $SO(16)$'s are broken to subgroups with $b \leq 0$.**

Stabilization

- The super no-scale models can be considered in a **cosmological scenario**.
- They are all **stable at early times, if finite temperature effects are taken into account**.
- At finite T ,

$$\mathcal{V}_{1\text{-loop}} \longrightarrow \text{free energy}$$

$$(\text{mass})^2 \longrightarrow T^2 + (\text{mass})^2$$

- As the Universe expands, T decreases :
 - **As long as $T > m_{3/2}$** : No tachyons \implies the models are **stable**.
 - **When T crosses $m_{3/2}$: Higgs phase transition** take place.

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$\mathcal{N} = 2 \rightarrow 0$ and $\mathcal{N} = 1 \rightarrow 0$ super no-scale models

- Consider $T^2 \times \frac{T^4}{\mathbb{Z}_2}$ or $\frac{T^2 \times T^2 \times T^2}{\mathbb{Z}_2 \times \mathbb{Z}_2}$ and implement the **stringy Scherk-Schwarz** mechanism along the first T^2 .

In general, the untwisted sector and twisted sectors contribute to $n_F - n_B$, except is one \mathbb{Z}_2 is freely acting (no fixed point) :

$$\mathcal{V}_{1\text{-loop}} = \frac{1}{2} \mathcal{V}_{1\text{-loop}}^{\mathcal{N}=4 \rightarrow 0} \quad \text{or} \quad \frac{1}{4} \mathcal{V}_{1\text{-loop}}^{\mathcal{N}=4 \rightarrow 0}$$

- Without free \mathbb{Z}_2 , there are large **Im T_1** corrections of the KK modes to

$$\frac{16 \pi^2}{g_{\text{YM}}^2(\mu)} = k \frac{16 \pi^2}{g_{\text{string}}^2} + b \log \frac{M_s^2}{\mu^2} + b \left(\frac{\pi}{3} \text{Im } T_1 - \log \text{Im } T_1 + \mathcal{O}(1) \right)$$

When $b < 0$, a **fine tuning** of g_{string} is a priori required to cancel it.

With one free \mathbb{Z}_2 , we have an underlying $\mathcal{N} = 4 \rightarrow \mathcal{N} = 2$

\implies **No Im T_1 term : No “Decompactification Problem”**

Conclusion

- The **Super No-Scale Models** are those which preserve the **flatness** of the effective potential at 1-loop (up to $\mathcal{O}(e^{-cM_s/m_{3/2}})$).
- ⇒ **Bosons - Fermions degeneracy at the massless level.**
- Their **quantum stability** is guaranteed if :
 - There are no Non-Asymptotically Free gauge groups ($b > 0$), in the $\mathcal{N} = 4 \rightarrow 0$ case.
 - Or simply if finite T is greater than $m_{3/2}$.
- When such a model is stable, it makes sense to decouple gravity to obtain MSSM-like models in flat space and let the electroweak radiative breaking stabilize $m_{3/2}$.
 - [Alvarez-Gaume, Polchinski, Wise] [Ibanez, Ross]
 - [Ellis, Nanopoulos, Tamvakis]
 - [Kounnas, Lahanas, Nanopoulos, Quiros]
 - [Kounnas, Zwirner, Pavel]
- **Question** : Is the effective potential at genus $g \geq 2$ still flat ? Or do we have to impose more constraints to guaranty the flatness condition ?